

Sheet 5: Solutions

Problem 1

We have

$$\omega_{\alpha\beta} = \pi_{\alpha}^{\mu} \pi_{\beta}^{\nu} (\nabla_{\mu} u_{\nu} - \nabla_{\nu} u_{\mu}) \quad (5.1.1)$$

$$\text{with } \pi_{\alpha}^{\mu} = \delta_{\alpha}^{\mu} - \frac{u^{\mu} u_{\alpha}}{c^2} \quad (5.1.2)$$

hence

$$\begin{aligned} \omega_{\alpha\beta} &= \left(\delta_{\alpha}^{\mu} - \frac{u^{\mu} u_{\alpha}}{c^2} \right) \nabla_{\mu} u_{\beta} \\ &\quad - \left(\delta_{\beta}^{\nu} - \frac{u^{\nu} u_{\beta}}{c^2} \right) \nabla_{\nu} u_{\alpha} \end{aligned}$$

$$= \nabla_{\alpha} u_{\beta} - \nabla_{\beta} u_{\alpha} - \frac{1}{c^2} (u_{\alpha} a_{\beta} - u_{\beta} a_{\alpha}) \quad (5.1.3)$$

$$\text{or } \omega = du^{\downarrow} - \frac{1}{c^2} u \wedge a$$

$$\Rightarrow du^{\downarrow} = \omega + \frac{1}{c^2} u \wedge a \quad (5.1.4)$$

$$u^{\downarrow} \wedge du^{\downarrow} = u \wedge \omega \quad (5.1.5)$$

$$\begin{aligned} i_u (u^{\downarrow} \wedge du^{\downarrow}) &= i_u (u \wedge \omega) \\ &= c^2 \omega - u \underbrace{i_u \omega}_{=0} \end{aligned}$$

Hence

$$\omega = \frac{1}{c^2} i_u (u^\dagger \wedge d u^\dagger) \quad (5.1.6)$$

Note: For any k -form Ω the map

$$\Omega \mapsto \frac{1}{c^2} i_u^\dagger (u^\dagger \wedge \Omega) \quad (5.1.7)$$

equals the projection \perp to u .

In components the proof is more involved.

Recall that the wedge-product of a p - and a q -form in components is

$$\lambda = \frac{1}{p!} \lambda_{a_1 \dots a_p} dx^{a_1} \wedge \dots \wedge dx^{a_p}$$

$$\eta = \frac{1}{q!} \eta_{a_1 \dots a_q} dx^{a_1} \wedge \dots \wedge dx^{a_q}$$

$$(\lambda \wedge \eta) = \frac{1}{(p+q)!} (\lambda \wedge \eta)_{a_1 \dots a_{p+q}} dx^{a_1} \wedge \dots \wedge dx^{a_{p+q}}$$

where

$$(\lambda \wedge \eta)_{a_1 \dots a_{p+q}} = \frac{(p+q)!}{p! q!} \lambda_{[a_1 \dots a_p} \eta_{a_{p+1} \dots a_{p+q}]}$$

Hence

$$(u^\dagger \wedge d u^\dagger)_{\alpha\beta\gamma} = \frac{(1+2)!}{1! 2!} u_{[\alpha} (\nabla_{\beta} u_{\gamma]} - \nabla_{\gamma} u_{\beta]) \quad (5.1.8)$$

$$(U^\dagger \wedge dU^\dagger)_{\alpha\beta\gamma}$$

$$= U_\alpha \nabla_\beta U_\gamma + U_\beta \nabla_\gamma U_\alpha + U_\gamma \nabla_\alpha U_\beta \\ - U_\beta \nabla_\alpha U_\gamma - U_\alpha \nabla_\gamma U_\beta - U_\gamma \nabla_\beta U_\alpha \quad (5.1.9)$$

$$[\text{i}_U (U^\dagger \wedge dU^\dagger)]_{\beta\gamma}$$

$$= U^\alpha (U^\dagger \wedge dU^\dagger)_{\alpha\beta\gamma}$$

$$= c^2 (\nabla_\beta U_\gamma - \nabla_\gamma U_\beta) - (U_\beta a_\gamma - U_\gamma a_\beta)$$

$$= c^2 \omega_{\beta\gamma} \quad (5.1.10)$$

↑ according to (5.1.3)

Hence once more

$$\frac{1}{c^2} \text{i}_U (U^\dagger \wedge dU^\dagger) = \omega. \quad (5.1.11)$$

Problem 2

$$g = c^2 dt \otimes dt - \delta_{ab} dx^a \otimes dx^b \quad (5.2.1)$$

$$K = \partial_t + \epsilon_{ab} \Omega^a X^b \partial_c \quad (5.2.2)$$

∂_t is itself a killing field. We have

$$\begin{aligned} L_K g &= -\delta_{ab} d(L_K X^a) \otimes dx^b \\ &\quad - \delta_{ab} dx^a \otimes d(L_K X^b) \end{aligned} \quad (5.2.3)$$

$$\text{But } L_K X^a = \epsilon_{mn}^a \Omega^m X^n$$

$$\leadsto d(L_K X^a) = \Omega^m \epsilon_{mn}^a dx^n \quad (5.2.4)$$

hence

$$\begin{aligned} L_K g &= -\delta_{ab} \Omega^m (\epsilon_{mn}^a dx^n \otimes dx^b \\ &\quad + \epsilon_{mn}^b dx^a \otimes dx^n) \\ &= -\Omega^m (\epsilon_{mnb} dx^n \otimes dx^b \\ &\quad + \epsilon_{mna} dx^a \otimes dx^n) \\ &= -\Omega^m (\epsilon_{mab} + \epsilon_{mba}) dx^a \otimes dx^b \\ &= 0 \end{aligned} \quad (5.2.5)$$

The corresponding 1-form to K is

$$K^\flat = g(K, \cdot) \quad (5.2.6)$$

$$\begin{aligned}
 K^\downarrow &= c^2 dt - \delta_{ab} \epsilon_{mn}^a \Omega^m X^n dX^b \\
 &= c^2 dt - \epsilon_{abc} \Omega^a X^b dX^c \quad (5.2.7)
 \end{aligned}$$

hence

$$\begin{aligned}
 g(K, K) &= K^\downarrow(K) = \\
 &= c^2 - \epsilon_{abc} \epsilon_{mn}^c \Omega^a X^b \Omega^m X^n \\
 &= c^2 - (\Omega \times \vec{X})^2 \\
 &= c^2 - \vec{\Omega}^2 \vec{X}^2 - (\vec{\Omega} \cdot \vec{X})^2 \\
 &= c^2 - \vec{\Omega}^2 \left(\vec{X} - \vec{n} (\vec{n} \cdot \vec{X}) \right)^2 \\
 &\quad \text{where } \vec{n} = \vec{\Omega} / \|\vec{\Omega}\| \\
 &= c^2 - \Omega^2 \vec{X}_\perp^2 \quad (5.2.8)
 \end{aligned}$$

For this to be ≥ 0 we must have

$$\|\vec{X}_\perp\| \leq c / \|\vec{\Omega}\| \quad (5.2.9)$$

Hence K is timelike in the space-time cylinder

$$U_K := \{(X^0, \vec{X}) \in M : \|\vec{X}_\perp\| \leq c / \|\vec{\Omega}\|\} \quad (5.2.10)$$

From (5.2.7) get

$$dk^\downarrow = -\epsilon_{abc} \Omega^a dx^b \wedge dx^c \quad (5.2.11)$$

$$K^\downarrow \wedge dk^\downarrow = -\epsilon_{abc} \Omega^a c^2 dt \wedge dx^b \wedge dx^c + \epsilon_{mnd} \Omega^m X^n \epsilon_{abc} \Omega^a dx^d \wedge dx^b \wedge dx^c \quad (5.2.12)$$

The second term can be rewritten using

$$dx^d \wedge dx^b \wedge dx^c = \epsilon^{dbc} dx^1 \wedge dx^2 \wedge dx^3 \quad (5.2.13)$$

$$\begin{aligned} 2^{\text{nd}} \text{ Term} &= \epsilon_{mnd} \underbrace{\epsilon_{abc} \epsilon^{dbc}}_{2\delta_{na}} \Omega^m \Omega^a X^n \\ &= 2 \epsilon_{mna} \Omega^m \Omega^a X^n = 0 \quad (5.2.14) \end{aligned}$$

Hence

$$K^\downarrow \wedge dk^\downarrow = -c \Omega^a \epsilon_{abc} dx^0 \wedge dx^a \wedge dx^b \quad (5.2.15)$$

Problem 3

$$g = c^2 dt \otimes dt - dr \otimes dr - r^2 d\varphi \otimes d\varphi \quad (5.3.1)$$

$$\varphi \mapsto \psi := \varphi - \Omega t \quad (5.3.2)$$

$$\begin{aligned} \leadsto g &= c^2 dt \otimes dt - dr \otimes dr \\ &\quad - r^2 (d\varphi + \Omega dt) \otimes (d\varphi + \Omega dt) \\ &= (c^2 - \Omega^2 r^2) dt \otimes dt - dr \otimes dr \\ &\quad - r^2 d\varphi \otimes d\varphi - r^2 \Omega (d\varphi \otimes dt + dt \otimes d\varphi) \end{aligned}$$

$$\begin{aligned} &= \left[1 - \left(\frac{\Omega r}{c} \right)^2 \right] \left[dx^0 - \frac{r^2 \Omega / c}{1 - \left(\frac{\Omega r}{c} \right)^2} d\varphi \right]^{\otimes 2} \\ &\quad - dr \otimes dr - r^2 \left[1 + \frac{(\Omega r / c)^2}{1 - \left(\frac{\Omega r}{c} \right)^2} \right] d\varphi \otimes d\varphi \end{aligned}$$

$$\begin{aligned} &= \left[1 - \left(\frac{\Omega r}{c} \right)^2 \right] (dx^0 + A)^{\otimes 2} - dr \otimes dr \\ &\quad - \frac{r^2}{1 - \left(\frac{\Omega r}{c} \right)^2} d\varphi \otimes d\varphi \quad (5.3.3) \end{aligned}$$

$$= \phi^2(r) \theta \otimes \theta - h \quad (5.3.4)$$

Where

$$\phi(\tau) = \left[1 - \left(\frac{\Omega \tau}{c} \right)^2 \right]^{1/2} \quad (5.3.5)$$

$$\Theta = dx^0 + A \quad (5.3.6)$$

$$A = \frac{\tau^2 \Omega / c}{1 - (\tau \Omega / c)^2} d\psi \quad (5.3.7)$$

$$h = dt \otimes dt + \frac{\tau^2 d\psi \otimes d\psi}{1 - (\tau \Omega / c)^2} \quad (5.3.8)$$

What 2-dim. geometry describes h ?

The circle of constant radius τ is given by curve $\gamma(\lambda) = (\tau(\lambda), \psi(\lambda))$:

$$\gamma(\lambda) = (\tau, \lambda), \quad \lambda \in [0, 2\pi] \quad (5.3.9)$$

$$\dot{\gamma} = (0, 1)$$

$$u(\tau) = \int_0^{2\pi} [h(\dot{\gamma}, \dot{\gamma})]^{1/2} d\lambda \quad (5.3.10)$$

$$= \frac{2\pi\tau}{\sqrt{1 - \left(\frac{\tau\Omega}{c}\right)^2}} \quad (5.3.11)$$

The radial distance along the curve

$$\gamma(\lambda) = (\lambda, 0), \quad \lambda \in [0, \tau]$$

$$\begin{aligned}
 R(\tau) &= \int_0^\tau (h(\dot{x}, \dot{x}))^{1/2} dx \\
 &= \int_0^\tau [h(\dot{x}/\tau, \dot{x}/\tau)] dx \\
 &= \int_0^\tau dx = \tau \quad (5.3.12)
 \end{aligned}$$

Hence circumference divided by radius is

$$\frac{U(\tau)}{R(\tau)} = \frac{2\pi}{\sqrt{1 - \left(\frac{\tau\Omega}{c}\right)^2}} > 2\pi \quad (5.3.13)$$

Can you visualize a 2-dim. surface embedded in \mathbb{R}^3 whose induced geometry has that behaviour and is rotationally symmetric? The answer must be no!, because this is impossible. The surface must be negatively curved, i.e. saddle-like, which cannot be embedded isometrically and be invariant under rotations at the same time.

Here is a calculation of the curvature tensor using Cartan structure equations.

We introduce orthonormal 1-forms

$$\theta^1 = dr \quad (5.3.14)$$

$$\theta^2 = \frac{r}{\sqrt{1 - \left(\frac{2M}{r}\right)^2}} d\psi \quad (5.3.15)$$

so that

$$h = \theta^1 \otimes \theta^1 + \theta^2 \otimes \theta^2 \quad (5.3.16)$$

Cartan's 1. structure equation is used to calculate the connection 1-form $\omega^a_b = \omega^a_{cb} \theta^c$ for the Levi-Civita connection

$$\boxed{d\theta^a + \omega^a_b \wedge \theta^b = 0} \quad (5.3.17)$$

$$\text{with } \omega_{ab} = -\omega_{ba} \quad (5.3.18)$$

$$\text{where } \omega_{ab} = \delta_{ac} \omega^c_b \quad (5.3.19)$$

$$\text{Set } a=1 \quad \theta^1 = dr \Rightarrow d\theta^1 = 0$$

$$\Rightarrow \omega^1_2 \wedge \theta^2 = 0$$

$$\Leftrightarrow \omega^1_2 \sim \theta^2 \quad (5.3.20)$$

$$\underline{\text{Set } a=2} \quad \theta^2 = f(r) d\varphi \quad (5.3.21)$$

$$\text{Where we set } f(r) := \frac{1}{\sqrt{1 - \left(\frac{\Omega r}{c}\right)^2}} \quad (5.3.22)$$

Then

$$\begin{aligned} d\theta^2 &= f'(r) dr \wedge d\varphi \\ &= (f'/f) \theta^1 \wedge \theta^2 \\ &= (-f'/f) \theta^2 \wedge \theta^1 \\ &= -\omega^2_1 \wedge \theta^1 \end{aligned} \quad (5.3.23)$$

$$\Leftrightarrow \omega^2_1 = (f'/f) \theta^2 + \sim \theta^1 \quad (5.3.24)$$

Now, since

$$\omega^2_1 = \omega_{21} = -\omega_{12} = -\omega^1_2 \quad (5.3.25)$$

we get from (5.3.20) and (5.3.24) together that

$$\begin{aligned} \omega^1_2 &= -\omega^2_1 = -\left(\frac{f'}{f}\right) \theta^2 \\ &= -f' d\varphi \end{aligned} \quad (5.3.26)$$

Cartan's 2. structure equation is used to calculate the curvature 2-form $\Omega^a_b = \frac{1}{2} R^a_{bcd} \theta^c \wedge \theta^d$ from the ω^a_b 's:

$$\boxed{\Omega^a_b = d\omega^a_b + \omega^a_c \wedge \omega^c_b} \quad (5.3.27)$$

In 2 dimensions, because $\omega^1_2 = -\omega^2_1$, is the only non-zero component of ω ($\omega^1_1 = \omega^2_2 = 0$) the second term $\sim \omega \wedge \omega$ vanishes. Hence

$$\begin{aligned} \Omega^1_2 &= d\omega^1_2 = -f'' dt \wedge d\varphi \\ &= -\frac{f''}{f} \theta^1 \wedge \theta^2 \end{aligned} \quad (5.3.28)$$

Therefore the only independent non-zero component of the Riemann-tensor with respect to or the normal frame is

$$R^1_{212} = -f''/f \quad (5.3.29)$$

Ricci-tensor is

$$\left. \begin{aligned} R_{11} &= R^a_{1a1} = R^2_{121} = R^1_{212} \\ &= R_{22} = -f''/f \end{aligned} \right\} (5.3.30)$$

Ricci - scalar

$$\begin{aligned}
 R &= h^{ab} R_{ab} = \delta^{ab} R_{ab} \\
 &= R_{11} + R_{22} \\
 &= -2 f'' / f
 \end{aligned} \tag{5.3.31}$$

In our case

$$\begin{aligned}
 f &= r \left[1 - \left(\frac{\Omega r}{c} \right)^2 \right]^{-1/2} \\
 f' &= \left[\dots \right]^{-1/2} - \frac{r}{2} \left[\dots \right]^{-3/2} (-2) \frac{r \Omega^2}{c^2} \\
 &= \left[\dots \right]^{-1/2} + \left[\dots \right]^{-3/2} \left(\frac{r \Omega}{c} \right)^2 \\
 &= \left[1 - \left(\frac{\Omega r}{c} \right)^2 \right]^{-3/2}
 \end{aligned} \tag{5.3.32}$$

$$\begin{aligned}
 f'' &= -\frac{3}{2} \left[\dots \right]^{-5/2} (-2) \frac{\Omega^2}{c^2} r \\
 &= 3 \left[1 - \left(\frac{\Omega r}{c} \right)^2 \right]^{-5/2} \left(\frac{\Omega}{c} \right)^2 r
 \end{aligned} \tag{5.3.33}$$

$$\Rightarrow f'' / f = 3 \left[\frac{\Omega / c}{1 - \left(\frac{\Omega r}{c} \right)^2} \right]^2 \tag{5.3.34}$$

Hence the Gaussian Curvature of the (r, ψ) -plane is

$$K = R_{1212} = -3 \left[\frac{\Omega/c}{1 - (\frac{\Omega r}{c})^2} \right]^2 \quad (5.3.31)$$

which is negative everywhere, smallest in modulus at the origin $r=0$

$$K(r=0) = -3 (\Omega/c)^2 \quad (5.3.32)$$

with monotonously increasing modulus for increasing r , diverging at $r = c/\Omega$ where the rotation of the frame becomes lightlike.

Problem 4

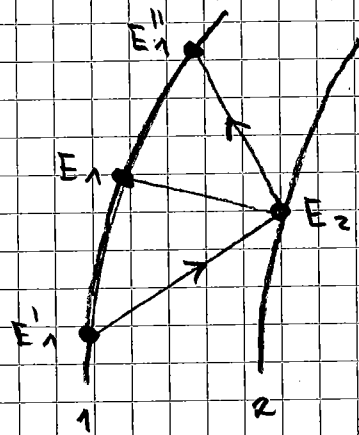
$$g = \phi^2 \Theta \otimes \Theta - h \quad (5.4.1)$$

$$i_K g = g(K, \cdot) = K^\flat = \phi^2 c \Theta \quad (5.4.2)$$

because $\Theta(K) = (dx^0 + A)(\partial/\partial t) = c$

Hence the kernel of K^\flat , which is the orthogonal complement of K , is identical to the kernel of Θ .

Consider two infinitesimally neighbouring orbits of K , called 1 and 2 and an



Einstein synchronization-procedure between them. That

is: Send a light signal from 1 to 2 where it

is reflected back to 1

Then the event E_1 on 1 that is declared to be "Einstein synchronous" with the reflection event E_2 on 2 is that half way (in terms of proper length) between the event E_1' of sending the signal from 1 and the event E_1'' of receiving the reflected light back at 1.

As is well known this "midpoint-condition" is equivalent to the condition that the vector connecting E_2 with E_1 is perpendicular to the world-line Z .

Hence a curve orthogonal to K , i.e. in the kernel of Θ , corresponds to a set of successively Einstein-simultaneous events. Hence, if γ is such a curve,

$$\left. \begin{aligned} \Theta(\dot{\gamma}) &= 0 \iff \\ dX^0(\dot{\gamma}) + A(\dot{\gamma}) &= 0 \end{aligned} \right\} (5.4.3)$$

i.e.

$$\int dX^0(\dot{\gamma}) d\lambda = \Delta X^0 = - \int_{\gamma} A \quad (5.4.4)$$

So the integral of A along the path γ equals minus the difference in the X^0 -coordinate for successive Einstein-synchronised events. But

$$A = \frac{(\tau\Omega/c)}{1 - (\tau\Omega/c)^2} \tau d\varphi \quad (5.4.5)$$

Integrated along a path whose spatial projection is a circle of radius τ ,

We get

$$\Delta t = - \frac{\Omega}{c^2} \frac{2\pi r^2}{1 - \left(\frac{\Omega r}{c}\right)^2} \quad (5.4.6)$$

or, since $\Delta \tau = \phi \Delta t$

$$\Delta \tau = - \frac{\Omega}{c^2} \frac{2\pi r^2}{\sqrt{1 - \left(\frac{\Omega r}{c}\right)^2}} \quad (5.4.7)$$

This lapse in Eigen time between the last and the first synchronised clock on a rotating frame in Minkowski space shows the failure of transitivity of Einstein Synchronisation if $K \wedge dK \neq 0$, i.e. if the world lines of the frame have non-zero vorticity. Note that the area of the spatial disk $r = \text{const}$ is

$$A(r) = \int_r \theta^1 \wedge \theta^2 = \int \frac{r}{\sqrt{1 - \left(\frac{\Omega r}{c}\right)^2}} dt \wedge d\varphi$$

$$= 2\pi \left(-\frac{c^2}{\Omega^2}\right) \left[1 - \left(\frac{\Omega r}{c}\right)^2\right]^{1/2} \Big|_0^r$$

$$= 2\pi \left(\frac{c}{\Omega}\right)^2 \left\{1 - \left[1 - \left(\frac{\Omega r}{c}\right)^2\right]^{1/2}\right\} \quad (5.4.8)$$

$$= \pi r^2 + \text{Terms} \sim \left(\frac{\Omega r}{c}\right)^{n \geq 2} \quad (5.4.9)$$

Note $A(r) > \pi r^2 \rightarrow \text{neg. curvature}$.

(5.4.7) and (5.4.9) give to leading order of Ω :

$$\Delta \tau = - \frac{\Omega}{c^2} 2 A(\tau) \quad (5.4.10)$$

Multiplied by the frequency $\nu = \frac{c}{\lambda}$ of an electromagnetic wave and multiplied by a factor of 2 for the difference between a clock - and anticlockwise synchronisation using light, this gives a phase difference of cycles

$$\Delta n = \frac{4 \Omega A}{\lambda c} \quad (5.4.10)$$

corresponding to a phase difference

$$\Delta \varphi = 2\pi \Delta n = \frac{8\pi \Omega A}{\lambda c} \quad (5.4.11)$$

This is just the phase shift observed in the Sagnac - Effect

Problem 5

$$g = \phi^2 c^2 dt \otimes dt - \bar{g}_{ab} dx^a \otimes dx^b \quad (5.5.1)$$

ϕ and \bar{g}_{ab} do not depend on t

Hence

$$\begin{aligned} \Gamma_{00}^0 &= \frac{1}{2} g^{0\lambda} (-g_{00,\lambda} + g_{\lambda 0,0} + g_{\phi\lambda,0}) \\ &= \frac{1}{2} g^{00} (-g_{00,0}) = 0 \quad \text{since } g^{0\lambda} = 0 \end{aligned} \quad (5.5.2)$$

$$\begin{aligned} \Gamma_{0a}^0 &= \Gamma_{a0}^0 = \frac{1}{2} g^{0\lambda} (-g_{\phi a,\lambda} + g_{\lambda 0,a} + g_{\phi\lambda,0}) \\ &= \frac{1}{2} g^{00} g_{00,a} = \phi_{,a} / \phi = [\ln \phi]_{,a} \end{aligned} \quad (5.5.3)$$

$$\begin{aligned} \Gamma_{00}^a &= \frac{1}{2} g^{a\lambda} (-g_{00,\lambda} + g_{\lambda 0,0} + g_{\phi\lambda,0}) \\ &= \frac{1}{2} g^{ab} (-g_{00,b}) \\ &= \frac{1}{2} \bar{g}^{ab} 2\phi\phi_{,b} = \bar{g}^{ab} \phi\phi_{,b} \end{aligned} \quad (5.5.4)$$

$$\begin{aligned} \Gamma_{ab}^0 &= \frac{1}{2} g^{0\lambda} (-g_{ab,\lambda} + g_{\lambda a,b} + g_{b\lambda,a}) \\ &= \frac{1}{2} g^{00} (-g_{\phi ab,0} + g_{\phi 0a,b} + g_{\phi b0,a}) \\ &= 0 \end{aligned} \quad (5.5.5)$$

$$\begin{aligned}
 \Gamma_{0b}^a &= \Gamma_{b0}^a \\
 &= \frac{1}{2} g^{a\lambda} (-g_{\cancel{0},\lambda} + g_{\cancel{0},0} + g_{0\lambda,b}) \\
 &= \frac{1}{2} g^{a0} g_{00,b} = 0 \quad (5.5.6)
 \end{aligned}$$

$$\begin{aligned}
 \Gamma_{bc}^a &= \frac{1}{2} g^{a\lambda} (-g_{bc,\lambda} + g_{\lambda a,b} + g_{b\lambda,a}) \\
 &= \frac{1}{2} g^{a\lambda} (-g_{bc,\lambda} + g_{\lambda a,b} + g_{b\lambda,a}) \\
 &= \frac{1}{2} g^{a\lambda} (-\bar{g}_{bc,\lambda} + \bar{g}_{b\lambda a} + \bar{g}_{ab,\lambda}) \\
 &= \bar{\Gamma}_{bc}^a \quad (5.5.7)
 \end{aligned}$$

Riemann tensor

$$\begin{aligned}
 R^{\alpha}_{\beta\mu\nu} &= \partial_{\mu} \Gamma_{\nu}^{\alpha\beta} - \partial_{\nu} \Gamma_{\mu}^{\alpha\beta} \\
 &\quad \Gamma_{\mu}^{\alpha} \times \Gamma_{\nu}^{\delta\beta} - \Gamma_{\nu}^{\alpha} \times \Gamma_{\mu}^{\delta\beta} \quad (5.5.8)
 \end{aligned}$$

Ricci - tensor

$$\begin{aligned}
 R_{\alpha\beta} &= \partial_{\lambda} \Gamma_{\alpha\beta}^{\lambda} - \partial_{\beta} \Gamma_{\alpha}^{\lambda\lambda} \\
 &\quad + \Gamma_{\lambda\delta}^{\lambda} \Gamma_{\alpha\beta}^{\delta} - \Gamma_{\alpha}^{\lambda\delta} \Gamma_{\beta\lambda}^{\delta} \quad (5.5.9)
 \end{aligned}$$

Calculation of R_{00}

$$R_{00} = \partial_\lambda \Gamma_{00}^\lambda - \partial_0 \Gamma_{0\lambda}^\lambda + \Gamma_{\lambda\gamma}^\lambda \Gamma_{00}^\gamma - \Gamma_{0\lambda}^\lambda \Gamma_{00}^\lambda \quad (5.5.10)$$

$$\partial_\lambda \Gamma_{00}^\lambda = \partial_a \Gamma_{00}^a = \partial_a (\bar{g}^{ab} \phi \phi_{,b})$$

$$\begin{aligned} \Gamma_{\lambda\gamma}^\lambda \Gamma_{00}^\gamma &= \Gamma_{\lambda a}^\lambda \Gamma_{00}^a \\ &= \Gamma_{\lambda a}^0 \Gamma_{00}^a + \Gamma_{0a}^0 \Gamma_{00}^a \\ &= \Gamma_{\lambda a}^0 \bar{g}^{ab} \phi \phi_{,b} \\ &\quad + (\phi_{,a} / \phi) \bar{g}^{ab} \phi \phi_{,b} \\ &= \Gamma_{\lambda a}^0 \bar{g}^{ab} \phi \phi_{,b} + \bar{g}^{ab} \phi_{,a} \phi_{,b} \end{aligned} \quad (5.5.11)$$

$$\begin{aligned} \Gamma_{0\lambda}^\lambda \Gamma_{00}^\lambda &= \Gamma_{0a}^0 \Gamma_{00}^a + \Gamma_{00}^a \Gamma_{0a}^0 \\ &= 2 \Gamma_{0a}^0 \Gamma_{00}^a = 2 \bar{g}^{ab} \phi_{,a} \phi_{,b} \end{aligned} \quad (5.5.12)$$

$$\begin{aligned} \Rightarrow R_{00} &= \partial_a (\bar{g}^{ab} \phi \phi_{,b}) \\ &\quad + \Gamma_{\lambda a}^0 \bar{g}^{ab} \phi \phi_{,b} \\ &\quad - \bar{g}^{ab} \phi_{,a} \phi_{,b} \end{aligned} \quad (5.5.13)$$

Introducing the covariant derivative on the $t = \text{const.}$ submanifolds with respect to \bar{g} :

$$\bar{\nabla}_a V^b = \partial_a V^b + \bar{\Gamma}_a{}^b{}_c V^c \quad (5.5.14)$$

applied to $V^a = \bar{g}^{ab} \phi \phi_{,b}$ we see that

$$\begin{aligned} R_{00} &= \bar{\nabla}_a (\bar{g}^{ab} \phi \phi_{,b}) - \bar{g}^{ab} \bar{\nabla}_a \phi \bar{\nabla}_b \phi \\ &= \phi \bar{g}^{ab} \bar{\nabla}_a \bar{\nabla}_b \phi \\ &= \phi \bar{\Delta} \phi \end{aligned} \quad (5.5.15)$$

where $\bar{\Delta} = \bar{g}^{ab} \bar{\nabla}_a \bar{\nabla}_b$

is the Laplacian w.r.t. \bar{g}

Further components of the Ricci-tensor are

$$\begin{aligned} R_{0a} = R_{a0} &= \partial_\lambda \Gamma_{a0}^\lambda - \partial_0 \Gamma_{a\lambda}^\lambda \\ &\quad + \Gamma_{\lambda\gamma}^\lambda \Gamma_{a0}^\gamma - \Gamma_{a\gamma}^\lambda \Gamma_{0\lambda}^\gamma \end{aligned} \quad (5.5.16)$$

$$\partial_\lambda \Gamma_{a0}^\lambda = \partial_\lambda \cancel{\Gamma_{a0}^\lambda} + \partial_0 \cancel{\Gamma_{a0}^0} = 0$$

$$\begin{aligned} \Gamma_{\lambda\gamma}^\lambda \Gamma_{a0}^\gamma &= \Gamma_{\lambda 0}^\lambda \Gamma_{a0}^0 + \Gamma_{\lambda c}^\lambda \Gamma_{a0}^c \\ &= \cancel{\Gamma_{00}^0} \Gamma_{a0}^0 + \cancel{\Gamma_{\lambda 0}^\lambda} \Gamma_{a0}^0 = 0 \end{aligned}$$

$$\begin{aligned}
 \Gamma_{a\lambda}^{\lambda} \Gamma_{0\lambda}^{\lambda} &= \Gamma_{a\lambda}^{\lambda} \Gamma_{0\lambda}^{\lambda} + \Gamma_{a\lambda}^0 \Gamma_{0\lambda}^{\lambda} \\
 &= \cancel{\Gamma_{a\lambda}^{\lambda} \Gamma_{0\lambda}^0} + \Gamma_{a\lambda}^{\lambda} \Gamma_{0\lambda}^c \\
 &\quad + \Gamma_{a\lambda}^0 \Gamma_{0\lambda}^0 + \cancel{\Gamma_{a\lambda}^0 \Gamma_{0\lambda}^c} \\
 &= 0
 \end{aligned}$$

Hence

$$R_{a0} = 0 \quad (5.5.17)$$

Finally

$$\begin{aligned}
 R_{ab} &= \partial_{\lambda} \Gamma_{ab}^{\lambda} - \partial_b \Gamma_{a\lambda}^{\lambda} \\
 &\quad + \Gamma_{\lambda\lambda}^{\lambda} \Gamma_{ab}^{\lambda} - \Gamma_{a\lambda}^{\lambda} \Gamma_{b\lambda}^{\lambda} \\
 &= \bar{R}_{ab} \\
 &\quad + \cancel{\partial_a \Gamma_{ab}^0} - \partial_b \Gamma_{a0}^0 \\
 &\quad + \cancel{\Gamma_{\lambda 0}^{\lambda} \Gamma_{ab}^0} + \Gamma_{0c}^0 \Gamma_{ab}^c + \cancel{\Gamma_{a0}^0 \Gamma_{ab}^0} \\
 &\quad - \cancel{\Gamma_{a\lambda}^{\lambda} \Gamma_{b\lambda}^0} - \cancel{\Gamma_{a\lambda}^0 \Gamma_{b0}^c} - \Gamma_{a0}^0 \Gamma_{b0}^0 \\
 &= \bar{R}_{ab} - \partial_b (\phi_{,a} / \phi) \\
 &\quad + (\phi_{,c} / \phi) \bar{\Gamma}_{ab}^c - \phi_{,a} \phi_{,b} / \phi^2 \\
 &= \bar{R}_{ab} - \frac{\bar{\nabla}_a \bar{\nabla}_b \phi}{\phi^2} \quad (5.5.18)
 \end{aligned}$$

Einstein's equations for $\Lambda = 0$, $T_{\alpha\beta} = 0$
are equivalent to

$$R_{\alpha\beta} = 0 \quad (5.5.19)$$

which in the static case hence become

$$\bar{\Delta} \phi = 0 \quad (5.5.20)$$

$$\bar{R}_{\alpha\beta} = \bar{\nabla}_\alpha \phi \bar{\nabla}_\beta \phi / \phi \quad (5.5.21)$$

If ϕ is regular everywhere and $\phi(\vec{x} \rightarrow \infty) = 1$, then $\phi \equiv 1$ since a harmonic function assumes its extrema on the boundary. Another proof is this: If ϕ is harmonic and asymptotically 1 then $\psi := \phi - 1$ is harmonic and asymptotically 0. Then on $\Sigma = \{x \in M : t = \text{const}\}$

$$\begin{aligned} & \int_{\Sigma} \bar{g}^{ab} \bar{\nabla}_a \psi \bar{\nabla}_b \psi \, d\mu_{\Sigma}(\bar{g}) \\ &= \int \bar{\nabla}_a (\bar{g}^{ab} \psi \bar{\nabla}_b \psi) \, d\mu_{\Sigma}(\bar{g}) \\ & \quad - \int \psi \bar{\Delta} \psi \, d\mu_{\Sigma}(\bar{g}) \end{aligned}$$

Graß \downarrow

$$= \int_{\partial\Sigma} \underbrace{(\psi \bar{\nabla}_b \psi)}_{\uparrow=3} n^b \, d\mathcal{O}_{\Sigma}(\bar{g}) = 0 \quad (5.5.22)$$

Hence

$$0 = \int \underbrace{\bar{g}^{ab} \bar{\nabla}_a \psi \bar{\nabla}_b \psi}_{\text{pos. def.}} d\mu_{\Sigma}(\bar{g})$$

$$\Leftrightarrow \bar{\nabla}_a \psi = 0 \Leftrightarrow \psi \equiv 0$$

$$\Leftrightarrow \phi \equiv 1. \quad 5.5.23$$

If ϕ is constant $\equiv 1$, from (5.5.21)

$$\Rightarrow \bar{\text{Ric}} = 0 \Rightarrow \bar{\text{Riem}} = 0$$

$$\Rightarrow \bar{g} \text{ is flat}$$

$$\Rightarrow \bar{g} = dx^0 \otimes dx^0 - d\vec{x} \otimes d\vec{x} \quad (5.5.24)$$

= Minkowski

Problem 6

$$g = \phi^2 (dx^0 \otimes dx^0 - \hat{g}_{ab} dx^a \otimes dx^b) \quad (5.6.1)$$

where ϕ and \hat{g}_{ab} only depend on \vec{x} .

$$L = \frac{1}{2} \phi^2 \left((\dot{z}^0)^2 - \hat{g}_{ab} \dot{z}^a \dot{z}^b \right) \quad (5.6.2)$$

Euler Lagrange Equation

$$\frac{\partial L}{\partial \dot{z}^0} = \phi^2 \dot{z}^0, \quad \frac{\partial L}{\partial z^0} = 0 \quad (5.6.3)$$

$$\Rightarrow \phi^2 \dot{z}^0 = k = \text{const} \quad (5.6.4)$$

$$\text{or } \phi^2 \left[\ddot{z}^0 + 2 \frac{\dot{\phi}}{\phi} \dot{z}^0 \right] = 0 \quad (5.6.5)$$

$$\frac{\partial L}{\partial \dot{z}^a} = -\phi^2 \hat{g}_{ab} \dot{z}^b \quad (5.6.7)$$

$$\begin{aligned} \frac{d}{dt} \frac{\partial L}{\partial \dot{z}^a} &= -\phi^2 \hat{g}_{ab} \ddot{z}^b - \phi^2 \hat{g}_{anb} \dot{z}^n \dot{z}^b \\ &\quad - 2\phi \dot{\phi} \hat{g}_{ab} \dot{z}^b \end{aligned} \quad (5.6.8)$$

where $\dot{\phi} = \phi_{,n} \dot{z}^n$

$$\begin{aligned} \frac{\partial L}{\partial z^a} &= \frac{\phi_{,a}}{\phi} \phi^2 \left((\dot{z}^0)^2 - \hat{g}_{ab} \dot{z}^a \dot{z}^b \right) \\ &\quad - \frac{1}{2} \phi^2 \hat{g}_{nba} \dot{z}^n \dot{z}^b \end{aligned} \quad (5.6.9)$$

Hence

$$\begin{aligned} \frac{d}{d\lambda} \left[\frac{\partial L}{\partial \dot{z}^a} \right] - \frac{\partial L}{\partial z^a} \\ = \phi^2 \hat{g}_{ab} \ddot{z}^b - \phi^2 \frac{1}{2} \left(-\hat{g}_{nb,ra} + \hat{g}_{an,rb} + \hat{g}_{ab,rn} \right) \dot{z}^n \dot{z}^b \\ - 2 \phi \dot{\phi} \hat{g}_{ab} \dot{z}^b - \sigma \phi_{,a} / \phi = 0 \end{aligned} \quad (5.6.10)$$

where we used

$$\phi^2 \left((\dot{z}^0)^2 - \hat{g}_{ab} \dot{z}^a \dot{z}^b \right) = \sigma \quad (5.6.11)$$

with $\sigma = \text{const}$.

This is equivalent to

$$\begin{aligned} \ddot{z}^a + \hat{\Gamma}_{bc}^a \dot{z}^b \dot{z}^c + 2 \frac{\dot{\phi}}{\phi} \dot{z}^a \\ = -\sigma \hat{g}^{ab} \frac{\phi_{,b}}{\phi^3} \end{aligned} \quad (5.6.12)$$

Written in terms of a new parameter $\lambda' = \gamma(\lambda)$ this becomes ($' = d/d\lambda'$)

$$\begin{aligned} \gamma'^2 \left[z''^a + \hat{\Gamma}_{bc}^a z'^b z'^c \right] + \dot{\gamma} z'^a \\ + 2 \left(\dot{\phi} / \phi \right) z'^a = -\sigma \hat{g}^{ab} \phi_{,b} / \phi^3 \end{aligned} \quad (5.6.13)$$

Choosing the new parameter s.t.

$$\begin{aligned} \ddot{\phi} \dot{z}^a + z \frac{\dot{\phi}}{\phi} \dot{z}^a \\ = \left(\ddot{\phi} + z \frac{\dot{\phi}}{\phi} \dot{\phi} / \phi \right) \dot{z}^a = 0 \end{aligned}$$

i.e. such that

$$\ddot{\phi} + z \frac{\dot{\phi}}{\phi} \dot{\phi} = 0 \quad (5.6.14)$$

From (5.6.5) we see that this equation is satisfied by

$$\phi(\lambda) = z^0(\lambda) \quad (5.6.15)$$

Hence we can choose $\phi(\lambda) = z^0(\lambda)$.

From (5.6.4), i.e. $\dot{\phi} \dot{z}^0 = k = \text{const}$, we get

$$\begin{aligned} \ddot{z}^a + \hat{\Gamma}_{bc}^a \dot{z}^b \dot{z}^c = - \frac{\sigma}{\kappa^2} \phi \hat{g}^{ab} \phi_{,b} \\ = - \frac{\sigma}{2\kappa^2} \hat{g}^{ab} \phi^2_{,b} \quad (5.6.16) \end{aligned}$$

or writing $z^0 = ct$

$$\begin{aligned} \frac{d^2 z^a}{dt^2} + \hat{\Gamma}_{bc}^a \frac{dz^b}{dt} \frac{dz^c}{dt} = - \frac{\sigma c^2}{2\kappa^2} \hat{g}^{ab} (\phi^2)_{,b} \\ = - C \hat{g}^{ab} \phi^2_{,b} \quad (5.6.17) \end{aligned}$$

For lightlike geodesics we have

$$\phi^2 \left(\left(\dot{z}^0 \right)^2 - \hat{g}_{ab} \dot{z}^a \dot{z}^b \right) = \sigma = 0 \quad (5.6.18)$$

Then

$$\frac{d^2 z^a}{dt^2} + \Gamma_{bc}^a \frac{dz^b}{dt} \frac{dz^c}{dt} = 0 \quad (5.6.19)$$

Hence we have proven that the spatial projection of a lightlike geodesic is a spacelike geodesic in the optical metric

$$\hat{g}_{ab} := \frac{-g_{ab}}{g_{00}} \quad (5.6.20)$$

with static time t as affine parameter proportional to optical arc length.

Note that \hat{g}_{ab} is a Riemannian (i.e. positive definite) metric. Hence the spatial path minimizes the optical arc length:

$$\delta \int \underbrace{\left[\hat{g}_{ab}(z(\lambda)) \dot{z}^a(\lambda) \dot{z}^b(\lambda) \right]^{1/2}}_{dt} d\lambda = 0 \quad (5.6.21)$$

\Leftrightarrow Fermat's principle of "quickest arrival"
with respect to static-time t .