Exercises for the lecture on Special Topics in GR & Relativistic Cosmology by DOMENICO GIULINI

Sheet 1

Problem 1

The Poisson equation for the Newtonian gravitational potential ϕ reads

$$\Delta \phi = 4\pi G \rho, \qquad (1)$$

where $\rho \ge 0$ is the matter density.

Give the most general solution for constant ρ if space is A) \mathbb{R}^3 and B) the round 3-sphere S³ with radius R.

Problem 2

Consider a set of N point-masses m_a , $a = 1, \cdots, N$, under the influence of their own gravitational attraction according to Newton's law. The 3N equations of motion are given by

$$m_{a}\ddot{\vec{x}}_{a}(t) = \sum_{\substack{b=1\\b\neq a}}^{N} Gm_{a}m_{b}\frac{\vec{x}_{b}(t) - \vec{x}_{a}(t)}{\|\vec{x}_{b}(t) - \vec{x}_{a}(t)\|^{3}}.$$
 (2)

Show that

$$\mathfrak{m}_{\mathfrak{a}}\ddot{\vec{x}}_{\mathfrak{a}}(t) = -\vec{\nabla}_{\mathfrak{a}}V\big(\vec{x}_{1}(t),\cdots,\vec{x}_{N}(t)\big)\,, \tag{3a}$$

where $\vec{\nabla}_{\mathfrak{a}}\coloneqq \partial/\partial\vec{x}_{\mathfrak{a}}$ and

$$V(\vec{x}_{1}, \cdots, \vec{x}_{N}) = -\frac{1}{2} \sum_{\substack{a,b=1\\a\neq b}}^{N} \frac{Gm_{a}m_{b}}{\|\vec{x}_{b} - \vec{x}_{a}\|}.$$
 (3b)

Show further that

$$\sum_{\alpha=1}^{N} \vec{x}_{\alpha} \cdot \vec{\nabla}_{\alpha} V = -V.$$
(4)

Problem 3

This is a continuation of the pervious problem.

Seek solutions of (3) of the form (called "homothetic motions")

$$\vec{x}_{a}(t) = a(t)\vec{y}_{a} \tag{5}$$

with the same non-negative function a(t) for all a and N time-independent vectors \vec{y}_a . Any N-tuple of vectors $(\vec{y}_1, \dots, \vec{y}_N)$ for which a solution to (2) exists is called a *central configuration*. The aim of this and the following problems is to discuss, as complete as possible here, the restrictions the equations of motion (3) imposes onto the function a(t) and upon the locations \vec{y}_a .

Show first that a(t) must satisfy a differential equation of the form

$$\frac{1}{2}\dot{a}^2 + \frac{C}{a} = E \tag{6}$$

where C and E are constants. Furthermore, show that the constant C is given by

$$C := \ddot{a}a^{2} = \frac{V(\vec{y}_{1}, \cdots, \vec{y}_{N})}{\sum_{a=1}^{N} m_{a} \|\vec{y}_{a}\|^{2}} = -\kappa < 0$$
(7)

and hence negative. The modulus of C is called κ .

Problem 4

Find solutions of (3) for negative, zero, and positive E.

Problem 5

Show that if $(\vec{y}_1, \dots, \vec{y}_N)$ is a central configuration, so is $(c\vec{y}_1, \dots, c\vec{y}_N)$, where $c \in \mathbb{R} - \{0\}$. Argue that therefore the search for central configurations may without loss of generality be be restricted to those on the (3N - 1)-dimensional ellipsoid

$$\sum_{\alpha=1}^{N} m_{\alpha} \|\vec{y}_{\alpha}\|^{2} = 1.$$
(8)

Show further that if $(\vec{y}_1, \dots, \vec{y}_N)$ is a central configuration, so is $(D\vec{y}_1, \dots, D\vec{y}_N)$, where $D \in SO(3)$ is a rotation matrix.

Problem 6

Show that the condition for $(\vec{y}_1, \dots, \vec{y}_N)$ being a central configuration is equivalent to (as in (7) we write $\kappa := -C$, so as to have $\kappa > 0$)

$$\kappa m_{a} \vec{y}_{a} + \sum_{\substack{b=1\\b \neq a}}^{N} G m_{a} m_{b} \frac{\vec{y}_{b} - \vec{y}_{a}}{\|\vec{y}_{b} - \vec{y}_{a}\|^{3}} = 0.$$
(9)

Show that this implies

$$\sum_{\alpha=1}^{N} \mathfrak{m}_{\alpha} \vec{y}_{\alpha} = \vec{0} \,. \tag{10}$$

Show the general identity (M = $\sum_{\alpha} m_{\alpha}$ denotes the total mass)

$$m_{a}\vec{y}_{a} = \frac{1}{M} \sum_{\substack{b=1\\b\neq a}}^{N} m_{a}m_{b}(\vec{y}_{a} - \vec{y}_{b}) + \frac{m_{a}}{M} \sum_{b=1}^{N} m_{b}\vec{y}_{b}.$$
 (11)

Use this and (10) to rewrite (9) as

$$\sum_{\substack{b=1\\b\neq a}}^{N} \vec{F}_{ab} = \vec{0}, \qquad (12a)$$

where

$$\vec{F}_{ab} = \mathfrak{m}_{a}\mathfrak{m}_{b}(\vec{y}_{a} - \vec{y}_{b})\left(\frac{\kappa}{M} - \frac{G}{r_{ab}^{3}}\right)$$
(12b)

and $r_{ab} := \|\vec{y}_a - \vec{y}_b\|$ denote the mutual distances.

Hence special central configurations are given if the N mass points can be arranged in such a way that all 2-particle distances are the same and equal to

$$r_{ab} = \left(\frac{MG}{\kappa}\right)^{1/3}.$$
 (13)

Note the remarkable fact that this holds independent of whether the masses making up M are equal or vastly different. What would possible configurations for N = 2, 3, 4 be? What about N = 5?

Another interpretation of equation (9) is by seeking the stationary points of the positive real-valued function $F(\vec{y}_1, \dots, \vec{y}_N) := -V(\vec{y}_1, \dots, \vec{y}_N)$ with constraints (Nebenbedingungen) that $\|\vec{y}_{\alpha}\| = R$ for all α . Mathematically the stationary points correspond to the lowest energy configuration of N positive charges placed on a 2-sphere of radius R. This lends some physical intuition to central configurations.