Exercises for the lecture on Foundations and Applications of Special Relativity von DOMENICO GIULINI

Sheet 12

Problem 1

A point particle of mass m and charge q moves along a future-oriented timelike curve $\tau \rightarrow z(\tau) \in M$, where we take the parameter τ to be the eigentime; hence $\dot{z}^2 = c^2$, where $\dot{z} := dz/d\tau$. The components of J^{α} of the electric four-current density and the components of the energy-momentum tensor are now distributions on M and given by

$$J^{\alpha}(\mathbf{x}) := qc \int d\tau \,\delta^{(4)} \big(\mathbf{x} - \mathbf{z}(\tau) \big) \,\dot{\mathbf{z}}^{\alpha}(\tau) \,, \tag{1}$$

$$\mathsf{T}^{\alpha\beta}(\mathbf{x}) := \mathfrak{m} \mathfrak{c} \int \mathrm{d}\tau \, \delta^{(4)} \big(\mathbf{x} - \mathbf{z}(\tau) \big) \, \dot{\mathbf{z}}^{\alpha}(\tau) \dot{\mathbf{z}}^{\beta}(\tau) \,. \tag{2}$$

- 1. Justify these expressions.
- 2. Show that the expression (1) could just as well have been written in terms of any other parameter λ instead of τ .
- 3. Prove that $\nabla_{\alpha} J^{\alpha} = 0$ (as a distribution).
- 4. Prove that $\nabla_{\alpha} T^{\alpha\beta} = 0$ (as a distribution) iff $\ddot{z} = 0$, i.e. if the curve is a straight line.

Problem 2

1. Prove that the relativistic Maxwell-Equations for a given source J^{α} are the Euler-Lagrange-Equations for the following action functional:

$$S_{f}[A] = \int d^{4}x \frac{1}{c} \left\{ -\frac{1}{4\mu_{0}} F_{\alpha\beta}(x) F^{\alpha\beta}(x) - A_{\alpha}(x) J^{\alpha}(x) \right\} , \qquad (3)$$

where the field to be varied is A, on which F depends via $F_{\alpha\beta} = \nabla_{\alpha}A_{\beta} - \nabla_{\beta}A_{\alpha}$. Hint: If \mathscr{L} denotes the integrand of the action functional (called the Lagrange density), the Euler-Lagrange-Equations are

$$\nabla_{\alpha} \left(\frac{\partial \mathscr{L}}{\partial (\nabla_{\alpha} A_{\beta})} \right) - \frac{\partial \mathscr{L}}{\partial A_{\beta}} = 0.$$
(4)

2. Why did we choose the overall sign and factor the way we did? (The Euler-Lagrange-Equations are insensitive to that.)

3. A relativistic particle of mass m is described by a future-oriented timelike worldline $\lambda \to z(\lambda)$, where now λ is an arbitrary parameter. As we know, its action is

$$S_{p}[z] = -mc \int d\lambda \, \sqrt{z'_{\alpha}(\lambda) z'^{\alpha}(\lambda)} \tag{5}$$

If the particle carries a charge q, it has a four-current density given by (1) (with λ replacing τ). Suppose the particle moves in an electromagnetic field represented by A, then its interaction with the field contributes a term according to (3). Show that the total action of the particle, now including its interaction with the field, is

$$S_{p+int} = \int d\lambda \left\{ -mc \sqrt{z'_{\alpha}(\lambda) z'^{\alpha}(\lambda)} - q A_{\alpha}(z(\lambda)) z'^{\alpha}(\lambda) \right\}$$
(6)

and that its Euler-Lagrange-Equations are

$$\mathfrak{m}\ddot{z}^{\alpha}(\tau) = q \, \mathsf{F}^{\alpha}{}_{\beta}(z(\tau)) \, \dot{z}^{\beta}(\tau) \,, \tag{7}$$

where a dot is again the derivative with respect to the eigentime parameter τ which – we recall – obeys $cd\tau = \sqrt{z'^{\alpha}(\lambda)z'_{\alpha}(\lambda)} d\lambda$.

Hint: In this case, if L denotes the integrand for the integral (6) (the Lagrange function), the Euler-Lagrange-Equations are

$$\frac{\mathrm{d}}{\mathrm{d}\lambda} \left(\frac{\partial L}{\partial z'^{\beta}} \right) - \frac{\partial L}{\partial z^{\beta}} = 0.$$
(8)

Problem 3

The Abraham-Lorentz-Equation in "ordinary" electrodynamics is meant to describe the spatial curve $t \mapsto z(t)$ of a slowly (i.e. $\|\dot{z}\|/c \ll 1$) moving point-particle of mass m and charge q under the action of an external electromagnetic field (**E**, **B**) and the particle's own radiation-reaction force. It reads:

$$\mathbf{m}\ddot{\mathbf{z}} = \mathbf{q}\left(\mathbf{E} + \dot{\mathbf{z}} \times \mathbf{B}\right) + \sigma \,\mathbf{m}\,\,\ddot{\mathbf{z}}\,,\tag{9}$$

where σ is a constant with the physical dimension of time:

$$\sigma = \frac{\mu_0 q^2}{6\pi cm} \,. \tag{10}$$

1. Show that σ can be rewritten as

$$\sigma = \frac{4}{3} \cdot \frac{q^2}{8\pi\varepsilon_0 mc^2} \cdot \frac{1}{c}$$
(11)

and use that form to interpret σ as a travel-time for light over a certain distance. What is that distance? What is its numerical value for the electron?

2. The relativistic generalisation of the the Abraham-Lorentz-Equation is the Lorentz-Dirac-Equation, which is *not* restricted to small velocities. It reads in components for the worldline $\tau \mapsto z^{\alpha}(\tau)$, parametrised by the eigentime τ ,

$$\mathfrak{m} \ddot{z}^{\alpha} = \mathfrak{q} \, \mathsf{F}^{\alpha}{}_{\beta} \dot{z}^{\beta} + \mathfrak{m} \sigma \left(\ddot{z}^{\alpha} + \frac{\ddot{z}^{\beta} \ddot{z}_{\beta}}{\mathfrak{c}^{2}} \dot{z}^{\alpha} \right) \,. \tag{12}$$

Here a dot now indicates the proper-time derivative. Argue that this is indeed the "obvious" generalisation of the Abraham-Lorentz equation (9). In particular, argue why the second factor in the bracket on the right-hand side is necessary and that without it the equation would certainly not capture the physically intended meaning and have almost no solutions. (Hint: Multiply the equation with \dot{z}_{α} and recall that a four-acceleration \ddot{z} is always perpendicular to the four-velocity \dot{z} .) What is the geometric interpretation of the sum the two terms in the bracket? (Hint: Compare Problem 4 on Sheet 10).

- 3. In Problem 5 of Sheet 9 we discussed the solution for the equation of motion for a point particle in a constant electric field, given that the particle was initially at rest. In Problem 5 (part 4) of Sheet 10 we identified this solution to be of constant acceleration. Prove that this particular solution also solves the Lorentz-Dirac equation (for the same constant electric field).
- 4. Without external field the Lorentz-Dirac-Equation van be written as

$$\dot{u}^{\alpha} = \sigma \Big(\ddot{u}^{\alpha} + \dot{u}^{\beta} \dot{u}_{\beta} \, u^{\alpha} \Big) \,, \tag{13}$$

where we set $u^{\alpha} := \dot{z}^{\alpha}/c$ so that $u^{\alpha}u_{\alpha} = 1$. It follows from this equation that any solution remains in the timelike plane spanned by the initial vectors u and \dot{u} . Let this plane be spanned by the orthonormal vectors e_0 (timelike) and e_1 (spacelike). Then the components u^0 and u^1 of the vector $u = u^0 e_0 + u^1 e_1$, which obey $(u^0)^2 - (u^1)^2 = 1$ may be written in terms of the rapidity $\rho(\tau)$ as follows

$$\mathfrak{u}^0 = \cosh(\rho) \quad \text{and} \quad \mathfrak{u}^1 = \sinh(\rho) \,.$$
 (14)

Show that (13) is now equivalent to the *linear* equation

$$\ddot{\rho} = \dot{\rho} / \sigma \,. \tag{15}$$

Use this to determine the most general solution for $\rho(\tau)$ and hence for $z(\tau)$. Show that in addition to the inertial motions (straight worldlines) you get accelerated solutions with exponentially growing rapidity.

Problem 4

This is problem 6 of sheet 11 reloaded.